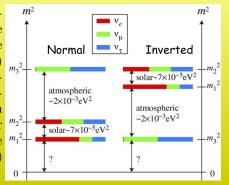
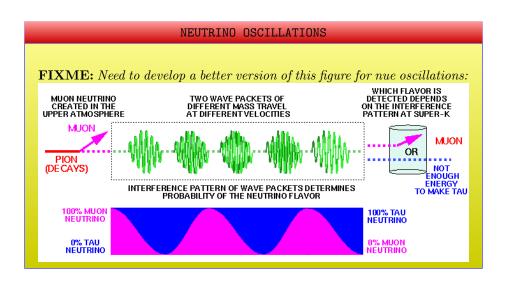
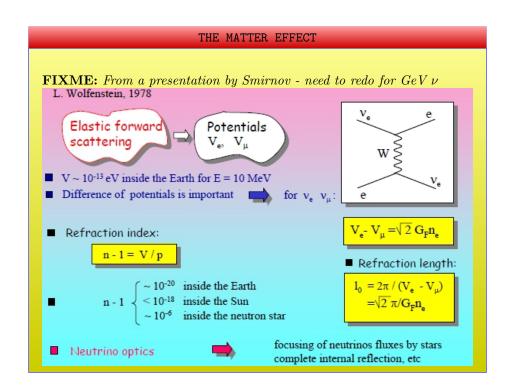
NEUTRINO MIXING

Neutrinos are weakly interacting elementary particles. There are three neutrino flavors which determine the lepton flavor (electron, muon, tau) produced in the weak charged current interaction of the neutrino. Neutrino flavors are not pure quantum eigenstates. Each neutrino flavor $(\nu_e, \nu_\mu, \nu_\tau)$ is an admixture of three quantum mass states (m_1, m_2, m_3) and vice-versa.



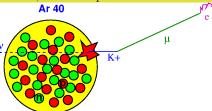




PROTON DECAY

FIXME: text taken verbatim from Wikipedia:

In the Standard Model, protons, are theoretically stable because baryon number (quark number) is conserved. Therefore, protons will not decay into other particles on their own, because they are the lightest (and therefore least energetic) baryon. Some beyond-the-Standard Model grand unified theories (GUTs) explicitly break the baryon number symmetry, allowing protons to decay via the Higgs particle, magnetic monopoles or new X bosons. Proton decay is one of the few observable effects of the various proposed GUTs. To date, all attempts to observe these events have failed.



SUPERNOVA BURST NEUTRINOS



1987A supernova remnant near the center. Composite of two public domain NASA images taken from the Hubble Space Telescope.

FIXME: taken verbatim from Wikipedia In 1966 Colgate and White calculated that neutrinos carry away most of the gravitational energy released by the collapse of massive stars, events now categorized as Type Ib and Ic and Type II supernovae. When such stars collapse, matter densities at the core becomes so high (1017 kg/m3) that the degeneracy of electrons is not enough to prevent protons and electrons from combining to form a neutron and an electron neutrino. A second and more important neutrino source is the thermal energy (100 billion kelvins) of the newly formed neutron core, which is dissipated via the formation of neutrino-antineutrino pairs of all flavors.